

Lecture #3 Hydrogenic Atoms (Chapter 13)

I. Applications of QM:

B. Rigid Rotor (RR) in a plane

C. Harmonic Oscillator

II. Experimental Observation of Line Spectra

III. QM Picture of Hydrogenic Atoms (13.1-13.3)

A. Energy Levels

B. Energy Functions (s,p,d)

I. Applications of QM

B. Rigid Rotor in a Plane

Model for rotational motion in which two bodies are separated by a rigid connector (good for modeling rotational energy of diatomic molecules assuming fixed bond length).

For a particle rotating on a ring of radius r in a plane: it's angular speed is $\omega = 2\pi\nu_{\text{rot}} = v/r$. Note $v =$ linear speed along tangent. The moment of inertia is $I = mr^2$, thus the angular momentum = $J = I\omega = mvr = p \times r$. The total **Energy = Kinetic Energy** = $I\omega^2/2 = J^2/2I = p^2r^2/2I$.

substituting the de Broglie relation (converts classical to QM): $p = h/\lambda$.

$$E = (hr/\lambda)^2/2I,$$

Imposing boundary conditions: An integral number of wavelengths must match circumference of path (justification seen via Fig. 12.27): $n\lambda = 2\pi r$, substituting this for λ in the above expression

$$\text{yields } E = (hn/2\pi)^2/2I = \frac{n^2\hbar^2}{2I} \text{ for } n = 0, \pm 1, \pm 2, \pm 3, \dots$$

$$\text{renaming } n \text{ as } m_l \text{ leads to a final expression: } E = \frac{m_l^2\hbar^2}{2I} \text{ for } m_l = 0, \pm 1, \pm 2, \pm 3, \dots$$

Correspondence Between Rectilinear and Angular Quantities

Linear	Angular
Mass (m)	Moment of Inertia ($I=mr^2$)
Speed ($v = 2\pi r\nu_{\text{rot}}$) cm/sec ($r = \text{cm/rad}$, $\nu_{\text{rot}} = \text{rev/sec}$)	Angular Speed ($\omega = 2\pi\nu_{\text{rot}} = v/r$) radians/sec
Momentum ($p=mv$)	Angular Momentum ($L=I\omega=mvr$) Kg $\text{m}^2 \text{s}^{-1}$ radians, usu. <i>radians</i> is excluded to have: Kg $\text{m}^2 \text{s}^{-1}$ ($h/2\pi$ units)
Kinetic Energy $mv^2/2 = p^2/(2m)$	Rotational Kinetic Energy $I\omega^2/2 = L^2/2I$

Exercise: For HCl rotating in a plane, if the moment of inertia (I) of HCl is $2.6 \times 10^{-47} \text{ Kg m}^2$, what energy in joules is needed to excite it from $m = 1$ to $m = 2$ level. What wavelength does this correspond to?

C. Harmonic Oscillator

Energy eigenvalue = $E_u = h\nu_0 (u + \frac{1}{2})$ $u = 0, 1, 2, 3, \dots$
 How would you write this equation in cm^{-1} ?

FUNDAMENTAL FREQUENCY $\ll k(\text{force constant})$

ν_0 = fundamental frequency in Hz, $\tilde{\nu}_0 = \frac{1}{2\pi c} \sqrt{\frac{k}{m}}$ = fundamental frequency in cm^{-1} ,

obtained from observing $u=0$ to $u=1$ transition from this we may find value of force constant k in N m^{-1} .

$$\Delta \bar{n}_{v=1} = \bar{n}_{v=1} - \bar{n}_{v=0} = \bar{n}_0(1 + \frac{1}{2}) - \bar{n}_0(0 + \frac{1}{2}) = \bar{n}_0$$

CAUTION WHEN COMPUTING m , make sure you divide molar masses by N_A to get "molecular mass"!

$\Delta E = h\nu =$ fundamental frequency. Given the observed fundamental freq \rightarrow calc k . For example, if for H^{35}Cl $n=0$ to $n=1$ has a fund freq of $2.9 \times 10^3 \text{ cm}^{-1}$, calculate k .

$$\bar{n}(\text{cm}^{-1}) = \frac{1}{2\pi c} \left(\frac{k}{m} \right)^{\frac{1}{2}} \left(u + \frac{1}{2} \right)$$

$$\Delta \bar{n}_{0 \rightarrow 1} = \frac{1}{2\pi c} \left(\frac{k}{m} \right)^{\frac{1}{2}} = 2.886 \times 10^3 \text{ cm}^{-1}$$

$$k = (2\pi \Delta \bar{n} c)^2 m$$

$$c = 2.998 \times 10^{10} \text{ cm s}^{-1}$$

$$m = \left[\frac{M_H M_{Cl}}{M_H + M_{Cl}} \right] \frac{1}{N_A} = \left[\frac{1.008 \text{ g mol}^{-1} \times 35.00 \text{ g mol}^{-1}}{1.008 \text{ g mol}^{-1} + 35.00 \text{ g mol}^{-1}} \right] \frac{1}{6.022 \times 10^{23} \text{ mol}^{-1}} \left(\frac{1 \text{ Kg}}{1000 \text{ g}} \right) = 1.627 \times 10^{-27} \text{ Kg}$$

$$k = 4.81 \times 10^2 \text{ N m}^{-1}$$

Purpose: Historical observation of H Line spectra. 2. Solving Schrodinger Equation to determine the energy levels (E) and wavefunctions (Y).

II. Experimental Observation of Line Spectra

When electric current passes through H_2 , it dissociates into electronically excited H atoms that emit radiation at discrete λ as they return to ground state.

The waveno. (λ & ν) fit the **Rydberg equation:** *Empirical*

$$\tilde{\nu} = R_H \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \quad n_1 = 1, 2, 3, \dots \quad n_2 = n_1 + 1, n_1 + 2, \dots \quad (1)$$

$n_1 = 1$ (Lyman), 2 (Balmer), 3 (Paschen), 4 (Brackett) and 5 (Pfund)

$R_H =$ Rydberg constant = 109677 cm^{-1}

What is the wavenumbers for a transition in the paschen series with $n_2 = 5$?

What is the frequency?

Bohr Frequency Condition: $\Delta E = h\nu$, Condition of resonance: E-diff between levels matches photon's E. These allowed E states are **stationary states**.

III. QM Picture of Hydrogenic Atoms: E Levels

Deriving Rydberg Expression from Bohr Postulates

Classical Rutherford Nuclear Model has a nucleus of charge Ze orbited by an electron, and predicts electron would fall to nucleus. But Bohr's stationary state hypothesis coupled with this presented a quantum mechanical picture with stable orbits. Solved the S.E. with Coulombic Potential to get energy eigenvalues and energy eigenfunctions (*stationary state*).

A. Energy Levels (Eigenvalues)

This was accomplished by setting boundary conditions in which the Coulombic force (V/r) balanced the centrifugal force and this result substituted into the Energy conservation equation: $E = \text{KE} + \text{Coulombic Potential } E$.

The Coulombic Potential Energy = $V = -Ze^2/(4\pi\epsilon_0 r)$ (2)

The resulting Energy Levels from boundary condition:

$E = -hcZ^2R/n^2$ where $R = 109677 \text{ cm}^{-1}$ (3) $n =$ principle quantum number. $n = 1, 2, 3, \dots$

$$\text{where } hcR = \frac{m e^4}{32\pi^2 \epsilon_0^2 \hbar^2} \quad (4)$$

$$m = \frac{m_e m_N}{m_e + m_N} \quad \text{Reduced mass} \quad (5) \quad \text{Since } m_N \gg m_e \text{ then } \mu \sim m_e.$$

Exercise E8.12

Shortest $\lambda =$ highest energy transition for $n = 3 \rightarrow n = \infty$, Thus waveno = $Z^2R(1/3^2)$ where $Z = 3$ for Li, thus waveno = 109677 cm^{-1} , and corresponding $\lambda = 91.2 \text{ nm}$.

Can compute **ionization energies** as energy from $n = 1$ to $n = \infty$: $I = E = -hcZ^2R$

HINT: It may be much easier to just use wavenumbers, i.e: $\tilde{\nu} = -Z^2R/n^2$ when you ultimately want to find λ .